

# An estimate of Earth-emitted night-light power captured by a satellite detector

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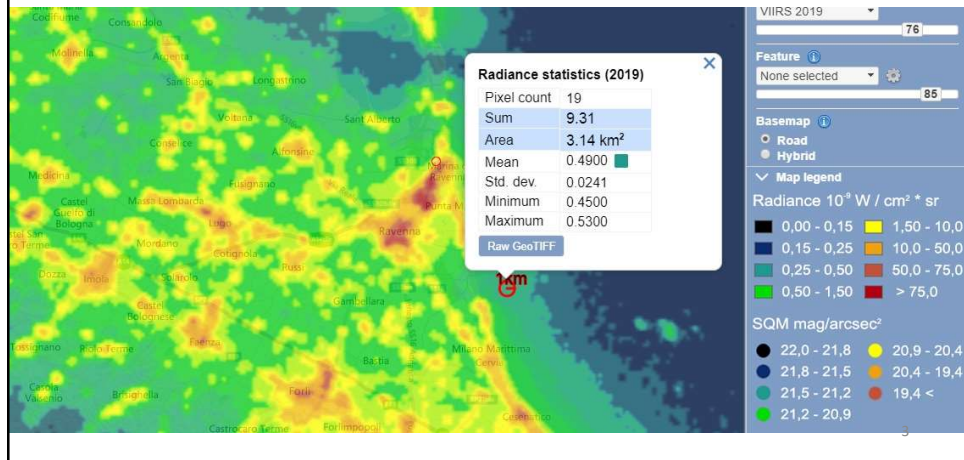
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## Contents

- Estimate the average night-light radiance ( $I_0$  in  $\text{W}/\text{m}^2/\text{sr}$ ) from publicly available data on the internet ([www.lightpollution.info](http://www.lightpollution.info)). Focusing on just 500 – 900 nm range.
- Assuming altitude ( $h$ ) and satellite detector diameter ( $d_s$ ), compute the total power the detector will capture as a function of the field-of-view (FOV).

From the website with light pollution data (VIIRS data)  
<https://www.lightpollutionmap.info/#zoom=9&lat=5533081&lon=1356725&ayers=B0TFFFFFFFTFFF>. One can determine the radiance in the 500-900nm wavelength range data. Using the built-in function on this website, I estimate the average radiance emitted is around  $I_0 = 10 \text{ uW/m}^2/\text{sr}$ .



## Compute the power at the satellite

- From VIIRS data, on the average, the light intensity (radiance) during the night is  $I_0 = 10 \text{ uW/m}^2/\text{sr}$ .
- The total power per  $\text{m}^2$  of Earth is given by,

$$I(\theta) = I_0 \cos \theta \quad d\Omega = \sin \theta \cdot d\theta d\phi$$

$$I_{total} = \int_0^{2\pi} \int_0^{\frac{\pi}{2}} I(\theta) \cos \theta \cdot d\Omega = I_0 \int_0^{2\pi} d\phi \int_0^{\frac{\pi}{2}} \cos \theta \cdot \sin \theta \cdot d\theta = \pi I_0$$

- $dA_E$  is the emitting surface on the Earth (FOV at the surface of Earth)
- $A_S$  is the receiving surface on the satellite (area is defined with  $A_S = d_S^2 \pi / 4$ , where  $d_S$  is the diameter of the satellite detector).

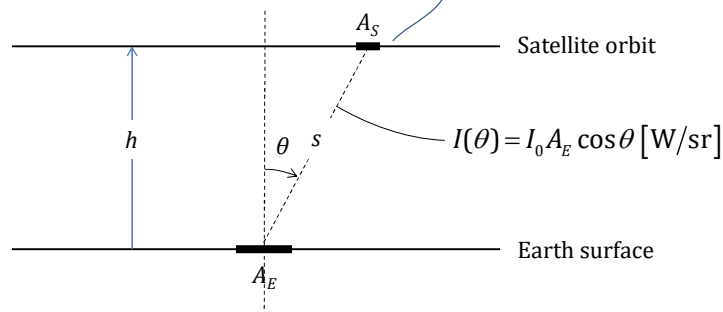
## Subtleties

- Surfaces  $dA_E$  and  $A_S$  are parallel to each other
- Satellite detector acceptance angle  $\Omega_S = 2\pi \cdot (1 - \cos\theta_{\max})$ .  
The etendue of the satellite detector is  $\Omega_S \cdot A_S$ .
- FOV is circle with diameter  $D = 2h \cdot \tan\theta_{\max}$ .

5

## Intensity received by $A_S$

$$\Omega_S(\theta) = \frac{A_S}{4\pi s^2} \cos\theta \text{ [sr]} = \frac{A_S}{4\pi h^2} \cos^3\theta \text{ [sr]}$$



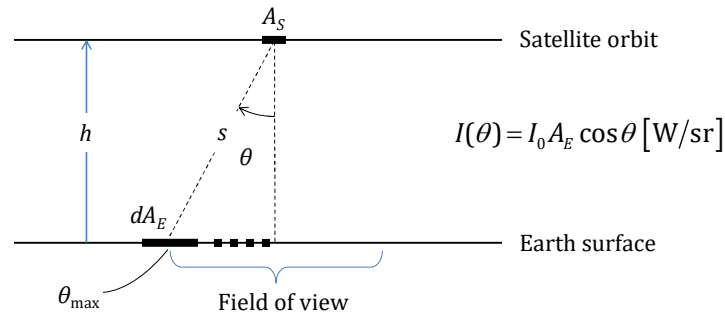
Intensity (in W) received by the satellite assuming that the receiving surface is parallel to Earth surface ( $A_S \parallel A_E$ )

$$I_S(\theta) = I(\theta) \cdot \Omega_S(\theta) = \frac{I_0 A_E A_S}{4\pi h^2} \cos^4\theta \text{ [W]}$$

6

## Sum $A_E$ over the field of view

Intensity (in W) received by the satellite assuming that the receiving surface is parallel to Earth surface ( $A_S \parallel A_E$ )



$$\frac{I_S(\theta)}{A_E} = \frac{I_0 A_S}{4\pi h^2} \cos^4 \theta \text{ [W/m}^2\text{]}$$

$$P_{TOT} = \int_0^{FOV} \left( \frac{I_S(\theta)}{A_E} \right) dA_E = \int_0^{FOV} \left( \frac{I_0 A_S}{4\pi h^2} \cos^4 \theta \right) dA_E \text{ [W]}$$

7

## Sum $A_E$ over the field of view

Now we sum the intensity over the entire FOV

$$\begin{aligned} dA_E &= 2\pi \cdot r \cdot dr = 2\pi \cdot \left( h \frac{\sin \theta}{\cos \theta} \right) \cdot \left( \frac{h d\theta}{\cos^2 \theta} \right) \\ &= 2\pi h^2 \cdot \frac{\sin \theta}{\cos^3 \theta} d\theta \end{aligned}$$

$$\begin{aligned} P_{TOT} &= \int_0^{FOV} \left( \frac{I_0 A_S}{4\pi h^2} \cos^4 \theta \right) dA_E = \int_0^{\theta_{\max}} \left( \frac{I_0 A_S}{4\pi h^2} \cos^4 \theta \right) \left( 2\pi h^2 \cdot \frac{\sin \theta}{\cos^3 \theta} \right) d\theta \\ &= \frac{I_0 A_S}{2} \int_0^{\theta_{\max}} \sin \theta \cos \theta d\theta \\ &= \frac{I_0 A_S}{4} (1 - \cos^2 \theta_{\max}) \end{aligned}$$

8

## Sum $A_E$ over the field of view

$$2h \tan \theta_{\max} = D_{\max}$$

$$\cos \theta_{\max} = \frac{1}{\sqrt{1 + \frac{D_{\max}^2}{4h^2}}}$$

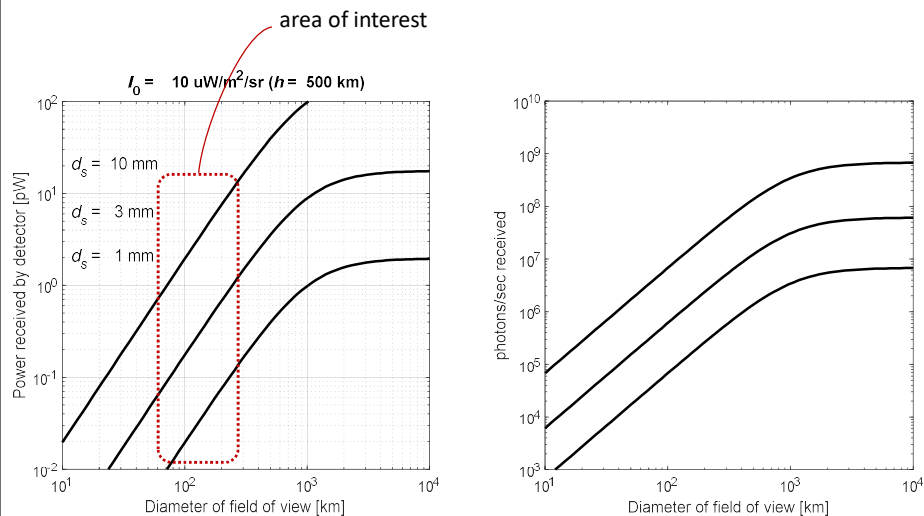
$$P_{TOT} = \frac{I_0 A_s}{4} (1 - \cos^2 \theta_{\max}) = \frac{I_0 d_s^2 \pi}{16} \left( \frac{D^2}{4h^2 + D^2} \right)$$

9

## Results

Note that peak emission from city centers is much higher.

\*Central Zagreb (120 m<sup>2</sup>) averages to 580 uW/m<sup>2</sup>/sr (VIIRS 2019)



## Future work

- Convert the power received into lumens to correlate to camera sensitivities.
- Estimate SNR for the TIA and detectors used